

# Bayesian Model Selection and Extrasolar Planet Detection

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This is a lovely description  
of how Bayesian Model  
Selection helps Extrasolar  
Planet Detection.

Data available on

(1) Amplitude of jitter

(2) Velocity offset

(3) Orbital period

(4) Velocity semi-amplitude

(5) Orbital eccentricity

(6) Orbital phase (2) One other variable

The catch is that even for moderately high dimensional  $\theta$  (as here) the integrals are not easy to calculate.

This is surprising since

(1)  $p(\theta | \text{data}, M_j)$  is "easy" to sample from by MCMC

(2)  $p(\theta | M_j), L(\theta | M_j)$  are usually easy to calculate and

$$(3) I = \frac{p(\theta | M_j) \cdot L(\theta | M_j)}{p(\theta | \text{data}, M_j)}$$

If we could calculate or estimate the posterior density (not just sample it) we could calculate or estimate I.

Suppose there are two models  $M_1, M_2$  with  $M_j$  postulating  $n_j$  planets.

Each model specifies a likelihood

$$L(\underline{\theta} | M_j) = \text{p.d.f. of data} \\ = p(\text{data} | \underline{\theta}, M_j)$$

The likelihood is assumed to be normal (eqn (1), p2).

The priors for different variables are given in Table 1, with a good discussion of reasons for choice.

The marginal likelihood of data given  $M_j$  is

$$I_j = p(\text{data} | M_j) = \int_{\underline{\theta}} L(\underline{\theta} | M_j) p(\underline{\theta} | M_j) d\underline{\theta}$$

Bayes Factor =  $p(\text{data} | M_2) / p(\text{data} | M_1)$

An estimate of  $I$  due to Raftery is (dropping  $M_j$  as index)

$$\left\{ \frac{1}{n} \sum_{i=1}^n \frac{1}{P(\theta_i) L(\theta_i)} \right\}^{-1} = \text{a Harmonic Mean}$$

where  $\theta_i$ 's are a sample from the posterior.

It doesn't ~~to~~ work because the mean square

$$\int \left\{ \frac{1}{P(\theta) L(\theta)} \right\}^2 P(\theta) L(\theta) d\theta$$

is typically infinite.

What about the new work of Raftery and Newton (Valencia 8)?

The paper discusses various other methods, e.g., ~~using~~ Laplace approximation, and a modification of Raftery's method due to Gelfand & Dey. Graphs and theoretical arguments show none of them perform well.

Two methods, Defensive Importance sampling and a new method of Berger — "Crazy Importance Sampling" — perform well. 6

Berger's method is based on the identity

$$I \stackrel{\text{def}}{=} \int p(\tilde{\theta}) L(\tilde{\theta}) d\tilde{\theta} = \frac{\int p(\tilde{\theta}) L(\tilde{\theta}) h(\tilde{\theta}) d\tilde{\theta}}{\int h(\tilde{\theta}) p(\tilde{\theta} | \text{data}) d\tilde{\theta}}$$

The trick is to sample  $\tilde{\theta}_i$ 's from  $h(\tilde{\theta})$  for the numerator and  $\bar{\theta}_i$ 's ~~to~~ from posterior for the denominator.

The estimate is

$$\frac{\sum p(\tilde{\theta}_i) L(\tilde{\theta}_i) / n'}{\sum p(\bar{\theta}_i) L(\bar{\theta}_i) / n}$$

(two imp samples)

I end this discussion with three new, possibly naive, methods.

1. The paper has a method of tracking multiple modes of  $p(\theta) L(\theta)$ . Given the location of the modes, it is possible to calculate Laplace Approximation around each mode and sum

2. Choose a suitable constant  $c$  which should be like the midpoint of the range of  $p(\theta) L(\theta)$ .

Let

$$I = \int p(\theta) L(\theta) d\theta + \int p(\theta) L(\theta) d\theta = I_1 + I_2$$

$$p(\theta) L(\theta) \leq c \quad p(\theta) L(\theta) > c$$

Estimate  $I_1$  by Berger's and  $I_2$  by Raftery's method.

Note that the problem of infinite variance of R's method will not arise here.

3. Let  $c$  be as in (2). Estimate  $I_3$  as in (2), i.e., by Raftery's method.

Now estimate

$$I_3 = \int p(\theta | \text{data}) d\theta$$
$$p(\theta) \mathcal{L}(\theta) \leq c$$

by  $\frac{1}{n} \sum \text{Indicator}\{p(\theta_i) \mathcal{L}(\theta_i) \leq c\}$   
where  $\theta_i$ 's are samples from  
the posterior by MCMC.

Now estimate  $I$  from  
the fact

$$I_3 = 1 - \frac{I_2}{I}$$

This is reminiscent of  
a method due to Chib.